Chapter 2: Pressure Distribution in a Fluid

Pressure and pressure gradient

In fluid statics, as well as in fluid dynamics, the forces acting on a portion of fluid (CV) bounded by a CS are of two kinds: body forces and surface forces.

Body Forces: act on the entire body of the fluid (force per unit volume).

Surface Forces: act at the CS and are due to the surrounding medium (force/unit area-stress).

In general the surface forces can be resolved into two components: one normal and one tangential to the surface. Considering a cubical fluid element, we see that the stress in a moving fluid comprises a 2\textsuperscript{nd} order tensor.
Since by definition, a fluid cannot withstand a shear stress without moving (deformation), a stationary fluid must necessarily be completely free of shear stress \((\sigma_{ij}=0, \ i \neq j)\). The only non-zero stress is the normal stress, which is referred to as pressure:

\[
\sigma_{ii} = -p
\]

\(\sigma_n = -p\), which is compressive, as it should be since fluid cannot withstand tension. (Sign convention based on the fact that \(p>0\) and in the direction of \(-n\)).

Or \(p_x = p_y = p_z = p_n = p\) (one value at a point, independent of direction; \(p\) is a scalar)

i.e. normal stress (pressure) is isotropic. This can be easily seen by considering the equilibrium of a wedge shaped fluid element.
\[ \sum F_x : -p_n \, dA \sin \alpha + p_x \, dA \sin \alpha = 0 \]
\[ p_n = p_x \]

\[ \sum F_z : -p_n \, dA \cos \alpha + p_z \, dA \cos \alpha - W = 0 \]

Where:
\[ W = \gamma V \]
\[ V = \Delta y - \Delta x \Delta z \]

\[ \Delta x = \Delta l \cos \alpha \quad \Delta z = \Delta l \sin \alpha \quad \Delta y \Delta l = dA \Rightarrow \Delta y = dA / dl \]

\[ W = \gamma dA \cos \alpha - dl \sin \alpha \]

\[ \Rightarrow -p_n \, dA \cos \alpha + p_z \, dA \cos \alpha - \gamma dA \cos \alpha - \frac{1}{2} dl \sin \alpha = 0 \]

\[ -p_n + p_z - \frac{\gamma}{2} dl \sin \alpha = 0 \]

\[ p_n = p_z \text{ for } dl \to 0 \text{ i.e. } p_n = p_x = p_y = p_z \]

Note: For a fluid in motion, the normal stress is different on each face and not equal to \( p \).
\[ \sigma_{xx} \neq \sigma_{yy} \neq \sigma_{zz} \neq -p \]

By convention \( p \) is defined as the average of the normal stresses
\[ p = -\frac{1}{3} (\sigma_{xx} + \sigma_{yy} + \sigma_{zz}) = -\frac{1}{3} \sigma_{ii} \]
The fluid element experiences a force on it as a result of the fluid pressure distribution if it varies spatially. Consider the net force in the x direction due to $p(x,t)$.

$$dF_{x_{net}} = pdydz - \left( p + \frac{\partial p}{\partial x} dx \right) dydz$$

$$= - \frac{\partial p}{\partial x} dxdydz$$

The result will be similar for $dF_y$ and $dF_z$; consequently, we conclude:

$$dF_{press} = \left[ - \frac{\partial p}{\partial x} \hat{i} - \frac{\partial p}{\partial y} \hat{j} - \frac{\partial p}{\partial z} \hat{k} \right] \Delta V$$

Or: $\underline{f} = -\nabla p$ force per unit volume due to $p(x,t)$.

Note: if $p=$constant, $\underline{f} = 0$. 
Equilibrium of a fluid element

Consider now a fluid element which is acted upon by both surface forces and a body force due to gravity

\[ \frac{dF_{\text{grav}}}{dV} = \rho g \quad \text{or} \quad f_{\text{grav}} = \rho g \]  
(per unit volume)

Application of Newton’s law yields: \( ma = \sum F \)

\[ \rho a = \sum f = f_{\text{body}} + f_{\text{surface}} \]  
per unit \( dV \)

\[ f_{\text{body}} = \rho g \quad \text{and} \quad g = -g \hat{k} \quad \Rightarrow \quad f_{\text{body}} = -\rho g \hat{k} \]

\[ f_{\text{surface}} = f_{\text{pressure}} + f_{\text{viscous}} \]

(includes \( f_{\text{viscous}} \), since in general \( \sigma_{ij} = -p\delta_{ij} + \tau_{ij} \))

\[ \begin{aligned} f_{\text{pressure}} &= -\nabla p \\
\mu \frac{\partial^2 V}{\partial x^2} + \frac{\partial^2 V}{\partial y^2} + \frac{\partial^2 V}{\partial z^2} &= \mu \nabla^2 V 
\end{aligned} \]

The viscous part

For \( \rho, \mu = \text{constant} \), the viscous force will have this form (chapter 4).

\[ \rho a = -\nabla p + \rho g + \mu \nabla^2 V \]

with \( a = \frac{\partial V}{\partial t} + V \cdot \nabla V \)
This is called the Navier-Stokes equation and will be discussed further in Chapter 4. Consider solving the N-S equation for $p$ when $a$ and $\mathbf{V}$ are known.

$$\nabla p = \rho(g - a) + \mu \nabla^2 \mathbf{V} = \mathbf{B}(\mathbf{x}, t)$$

This is simply a first order PDE for $p$ and can be solved readily. For the general case ($\mathbf{V}$ and $p$ unknown), one must solve the NS and continuity equations, which is a formidable task since the NS equations are a system of 2\textsuperscript{nd} order nonlinear PDEs.

We now consider the following special cases:

1) Hydrostatics ($a = \mathbf{V} = 0$)

2) Rigid body translation or rotation ($\nabla^2 \mathbf{V} = 0$)

3) Irrotational motion ($\nabla \times \mathbf{V} = 0$)

$$\nabla \times (\nabla \times \mathbf{b}) = \nabla(\nabla \cdot \mathbf{b}) - \nabla^2 \mathbf{b}$$

\textit{vector identity} \hspace{1cm} \text{For vector } \mathbf{b} = \mathbf{V}

if $\rho = \text{constant}$

$$\nabla \times \mathbf{V} = 0 \Rightarrow \nabla^2 \mathbf{V} = 0 \Rightarrow \text{Euler equation} \Rightarrow \int \Rightarrow \text{Bernoulli equation}$$

also,

$$\nabla \times \mathbf{V} = 0 \Rightarrow \mathbf{V} = \nabla \varphi \; \text{ & if } \rho = \text{const. } \Rightarrow \nabla^2 \varphi = 0$$
Case (1) Hydrostatic Pressure Distribution

\[ \nabla p = \rho g = -\rho g k \quad z \uparrow \quad \downarrow g \]

i.e. \( \frac{\partial p}{\partial x} = \frac{\partial p}{\partial y} = 0 \) and \( \frac{\partial p}{\partial z} = -\rho g \quad dp = -\rho g dz \)

or \( p_2 - p_1 = -\int_{1}^{2} \rho g dz = -g \int_{1}^{2} \rho(z)dz \)

\[ g = g_0 \left( \frac{r}{r_0} \right)^2 \]
\[ \approx \text{constant near earth's surface } r_0 \]

liquids \( \Rightarrow \rho = \text{constant (for one liquid)} \)
\( p = -\rho gz + \text{constant} \)

gases \( \Rightarrow \rho = \rho(p,t) \) which is known from the equation of state: \( p = \rho RT \Rightarrow \rho = p/RT \)

\[ \frac{dp}{p} = -\frac{g}{RT(z)} \quad \text{which can be integrated if } T=T(z) \text{ is} \]
\[ \text{known as it is for the atmosphere.} \]
Manometry

Manometers are devices that use liquid columns for measuring differences in pressure. A general procedure may be followed in working all manometer problems:

1.) Start at one end (or a meniscus if the circuit is continuous) and write the pressure there in an appropriate unit or symbol if it is unknown.

2.) Add to this the change in pressure (in the same unit) from one meniscus to the next (plus if the next meniscus is lower, minus if higher).

3.) Continue until the other end of the gage (or starting meniscus) is reached and equate the expression to the pressure at that point, known or unknown.
Hydrostatic forces on plane surfaces

The force on a body due to a pressure distribution is:

\[ F = - \int_{A} p n \, dA \]

where for a plane surface \( n = \text{constant} \) and we need only consider \( |F| \) noting that its direction is always towards the surface:

\[ |F| = \int_{A} p \, dA. \]

Consider a plane surface \( \overline{AB} \) entirely submerged in a liquid such that the plane of the surface intersects the free-surface with an angle \( \alpha \). The centroid of the surface is denoted \((\overline{x}, \overline{y})\).

\[ F = \gamma \sin \alpha \, \overline{y}A = \overline{p}A \]

Where \( \overline{p} \) is the pressure at the centroid.
To find the line of action of the force which we call the center of pressure \((x_{cp}, y_{cp})\) we equate the moment of the resultant force to that of the distributed force about any arbitrary axis.

\[
y_{cp}F = \int_A ydF = \gamma \sin \alpha \int_A y^2 dA \quad \text{Note: } dF = \gamma y \sin \alpha dA
\]

\[
\int_A y^2 dA = I_o \rightarrow \text{moment of Inertia about } O - O
\]

\[
= \bar{y}^2 A + \bar{I}
\]

\(\bar{I} = \text{moment of inertia WRT horizontal centroidal axis}\)

\[
\Rightarrow F = \bar{p}A = \gamma \sin \alpha \bar{y}A
\]

\[
\Rightarrow y_{cp} \gamma \sin \alpha \bar{y}A = \gamma \sin \alpha \left( \bar{y}^2 A + \bar{I} \right)
\]

\[
\Rightarrow y_{cp} = y + \frac{\bar{I}}{\bar{y}A}
\]

and similarly for \(x_{cp}\)

\[
x_{cp}F = \int_A xdF \quad \text{where} \quad \bar{I}_{xy} = \text{product of inertia}
\]

\[
\bar{I}_{xy} = \bar{I}_{xy} + \bar{x} \bar{y}A
\]

Note that the coordinate system in the text has its origin at the centroid and is related to the one just used by:

\[
x_{text} = x - \bar{x} \quad \text{and} \quad y_{text} = -(y - \bar{y})
\]
Hydrostatic Forces on Curved Surfaces

In general, 

\[ F = - \int_{A} p \mathbf{n} \cdot dA \]

Horizontal Components:

\[ F_x = F \cdot \hat{i} = - \int_{A} p \mathbf{n} \cdot \hat{i} \frac{dA}{dA_x} \]

\[ F_y = - \int_{A_y} p \ dA_y \]

\( dA_x \) = projection of \( \mathbf{n} \) \( dA \) onto a plane perpendicular to \( x \) direction

\( dA_y \) = projection of \( \mathbf{n} \) \( dA \) onto a plane perpendicular to \( y \) direction

The horizontal component of force acting on a curved surface is equal to the force acting on a vertical projection of that surface including both magnitude and line of action and can be determined by the methods developed for plane surfaces.

\[ F_z = - \int_{A_z} p \mathbf{n} \cdot \hat{k} \ dA = - \int_{A_z} p \ dA_z = \gamma \int_{A_z} h \ dA_z = \gamma \forall \]

Where \( h \) is the depth to any element area \( dA \) of the surface. The vertical component of force acting on a curved surface is equal to the net weight of the total column of fluid directly above the curved surface and has a line of action through the centroid of the fluid volume.
Example  Drum Gate

\[ h = R - R \cos \theta = R (1 - \cos \theta) \]

\[ p = \gamma h = \gamma R (1 - \cos \theta) \]

\[ \bar{n} = -\sin \theta \hat{i} + \cos \theta \hat{k} \]

\[ dA = l R d\theta \]

\[ F = -\int_{0}^{\pi} \gamma R (1 - \cos \theta) \left( -\sin \theta \hat{i} + \cos \theta \hat{k} \right) l R d\theta \]

\[ F \hat{i} = F_x = \gamma l R^2 \int_{0}^{\pi} (1 - \cos \theta) \sin \theta d\theta \]

\[ = \gamma l R^2 \left( -\cos \theta \bigg|_0^\pi + \frac{1}{4} \cos 2\theta \bigg|_0^\pi \right) = 2\gamma l R^2 \]

\[ = \frac{\gamma R^2 l}{p \bar{A}} \quad \Rightarrow \quad \text{Same force as that on projection of gate onto vertical plane perpendicular direction} \]
\[ F_z = -\gamma l R^2 \int_0^\pi (1 - \cos \theta) \cos \theta d\theta \]

\[ = -\gamma l R^2 \left( \sin \theta - \frac{\theta}{2} - \frac{1}{4} \sin 2\theta \right)_0^\pi \]

\[ = -\gamma l R^2 \frac{\pi}{2} = \gamma l \left( \frac{\pi R^2}{2} \right) = \gamma \forall \]

Net weight of water above curved surface

Another approach:

\[ F_1 = \gamma l \left[ R^2 - \frac{1}{4} \pi R^2 \right] \]

\[ = \gamma l R^2 \left[ 1 - \frac{1}{4} \pi \right] \]

\[ F_2 = \gamma l \frac{\pi R^2}{2} + F_1 \]

\[ F = F_2 - F_1 = \frac{\gamma l \pi R^2}{2} \]
Hydrostatic Forces in Layered Fluids
See textbook 2.7

Buoyancy and Stability

Archimedes Principle

\[ F_B = F_{V(2)} - F_{V(1)} \]

= fluid weight above \(2_{ABC} \) – fluid weight above \(1_{ADC} \)

= weight of fluid equivalent to the body volume

In general, \( F_B = \rho g \forall \) (\( \forall \) = submerged volume).

The line of action is through the centroid of the displaced volume, which is called the center of buoyancy.
Example: Floating body in heave motion

Weight of the block $W = \rho_b L b h g = m g = \gamma \forall_0$ where $\forall_0$ is displaced water volume by the block and $\gamma$ is the specific weight of the liquid.

$W = B \Rightarrow \rho_b L b h g = \rho_w L b d g \Rightarrow d = \frac{\rho_b}{\rho_w} h = S_b h$

$\rho_b = \rho_w : d = h$

$\rho_b > \rho_w : d > h$ sink

$\rho_b < \rho_w : d < h$ floating

Instantaneous displaced water volume:

$\forall = \forall_0 - y A_{wp}$

$\sum F_v = m \ddot{y} = B - W = \gamma \forall - \gamma \forall_0$

$= -\gamma A_{wp} y$
\[ m\ddot{y} + \gamma A_{wp} y = 0 \]

\[ \dot{y} + \frac{\gamma A_{wp}}{m} \dot{y} = 0 \]

\[ y = A \cos \omega_{n} t + B \sin \omega_{n} t \]

Use initial condition \((t = 0, \ y = y_0, \ \dot{y} = \dot{y}_0)\) to determine \(A\) and \(B\):

\[ y = y_0 \cos \omega_{n} t + \frac{\dot{y}_0}{\omega_{n}} \sin \omega_{n} t \]

Where

\[ \omega_{n} = \sqrt{\frac{\gamma A_{wp}}{m}} \]

\[ T = \frac{2\pi}{\omega} = 2\pi \sqrt{\frac{m}{\gamma A_{wp}}} \]  

Spar Buoy

T is tuned to decrease response to ambient waves: we can increase \(T\) by increasing block mass \(m\) and/or decreasing waterline area \(A_{wp}\).
Stability of Immersed and Floating Bodies

Here we’ll consider transverse stability. In actual applications both transverse and longitudinal stability are important.

Immersed Bodies

\[
\sum \mathbf{F} = 0 \quad \text{and} \quad \sum \mathbf{M} = 0
\]

\(\sum \mathbf{M} = 0\) requires that the centers of gravity and buoyancy coincide, i.e., \(C = G\) and body is neutrally stable.

If \(C\) is above \(G\), then the body is stable (righting moment when heeled).

If \(G\) is above \(C\), then the body is unstable (heeling moment when heeled).

**FIGURE 3.15**

Conditions of stability for immersed bodies.
(a) Stable. (b) Neutral. (c) Unstable.
Floating Bodies

For a floating body the situation is slightly more complicated since the center of buoyancy will generally shift when the body is rotated depending upon the shape of the body and the position in which it is floating.

The center of buoyancy (centroid of the displaced volume) shifts laterally to the right for the case shown because part of the original buoyant volume AOB is transferred to a new buoyant volume EOD.

The point of intersection of the lines of action of the buoyant force before and after heel is called the metacenter M and the distance GM is called the metacentric height. If GM is positive, that is, if M is above G, then the ship is stable; however, if GM is negative, the ship is unstable.
\( \alpha = \) small heel angle

\( x = CC' = \) lateral displacement of \( C \)

\( C = \) center of buoyancy
  
  i.e., centroid of displaced volume \( V \)

Solve for GM: find \( \bar{x} \) using

(1) basic definition for centroid of \( V \); and

(2) trigonometry

(1) Basic definition of centroid of volume \( V \)

\[
\bar{x}V = \int x dV = \sum x_i \Delta V_i \quad \text{moment about centerplane}
\]

\[
\bar{x}V = \text{moment } V \text{ before heel} - \text{moment of } V_{AOB} + \text{moment of } V_{EOD}
\]

\( = 0 \) due to symmetry of original \( V \) about \( y \) axis

i.e., ship centerplane

\[
\bar{x}V = - \int_{AOB} (-x) dV + \int_{EOD} x dV \quad \tan \alpha = y/x
\]

\[ dV = y dA = x \tan \alpha \ dA \]

\[
\bar{x}V = \int_{AOB} x^2 \tan \alpha dA + \int_{EOD} x^2 \tan \alpha dA
\]
\[ \bar{x}V = \tan \alpha \int x^2 dA \]

- **Ship waterplane area**
- **Moment of inertia of ship waterplane about z axis O-O; i.e., \( I_{oo} \)**

\[ I_{oo} = \text{moment of inertia of waterplane area about centerplane axis} \]

(2) Trigonometry

\[ \bar{x}V = \tan \alpha I_{oo} \]

\[ CC' = \bar{x} = \frac{\tan \alpha I_{oo}}{V} = CM \tan \alpha \]

\[ CM = \frac{I_{oo}}{V} \]

\[ GM = CM - CG \]

\[ GM = \frac{I_{oo}}{V} - CG \]

- **GM > 0** Stable
- **GM < 0** Unstable
**Roll:** The rotation of a ship about the longitudinal axis through the center of gravity.

Consider symmetrical ship heeled to a very small angle $\theta$. Solve for the subsequent motion due only to hydrostatic and gravitational forces.

\[
F_b = \left( \cos \hat{\theta}j - \sin \hat{\theta} \right) \rho g \forall \quad (\rho g \forall = \Delta)
\]

\[
M_g = r \times F_b
\]

\[
M_g = \left( -G C\hat{j} + C C'\hat{i} \right) \times \Delta \left( \cos \hat{\theta}j - \sin \hat{\theta} \right)
\]

\[
= \left( -G C \sin \theta + C C' \cos \theta \right) \Delta \hat{k}
\]

\[
= \left( -G C + C M \right) \sin \theta \Delta \hat{k}
\]

\[
= G M \sin \theta \Delta \hat{k}
\]

Note: recall that $M_o = |\overrightarrow{F}| \cdot d$, where $d$ is the perpendicular distance from $O$ to the line of action of $\overrightarrow{F}$.

\[
M_o = G Z \Delta
\]

\[
= G M \sin \theta \Delta
\]
\[ \sum M = -I \ddot{\theta} \]

I = mass moment of inertia about long axis through G  
\( \ddot{\theta} = \) angular acceleration

\[ I \ddot{\theta} + \Delta GM \sin \theta = 0 \]

for small \( \theta \):  
\[ \ddot{\theta} + \frac{\Delta GM}{I} \theta = 0 \]

\[ \frac{\Delta GM}{I} = \frac{\rho g \nabla GM}{I} = \frac{mgGM}{I} \]

\[ k = \sqrt{\frac{I}{m}} \text{ definition of radius of gyration} \]

\[ k^2 = \frac{I}{m} \quad mk^2 = I \quad \frac{\Delta GM}{I} = \frac{gGM}{k^2} \]

The solution to this equation is,

\[ \theta(t) = \theta_o \cos \omega_n t + \frac{\dot{\theta}}{\omega_n} \sin \omega_n t \]

0 for no initial velocity

where \( \theta_o = \) the initial heel angle

\( \omega_n = \) natural frequency

\[ = \sqrt{\frac{gGM}{k^2}} = \frac{\sqrt{gGM}}{k} \]
Simple (undamped) harmonic oscillation:

The period of the motion is

\[ T = \frac{2\pi}{\omega_n} \quad T = \frac{2\pi k}{\sqrt{gGM}} \]

Note that large \( GM \) decreases the period of roll, which would make for an uncomfortable boat ride (high frequency oscillation).

Earlier we found that \( GM \) should be positive if a ship is to have transverse stability and, generally speaking, the stability is increased for larger positive \( GM \). However, the present example shows that one encounters a “design tradeoff” since large \( GM \) decreases the period of roll, which makes for an uncomfortable ride.
Parametric Roll:

The periodicity of the encounter wave causes variations of the metacentric height i.e. $GM = GM(t)$. Therefore:

$$I \ddot{\theta} + \Delta GM(t) \theta = 0$$

Assuming $GM(t) = GM_0 + GM_1 \cos(\omega t)$:

$$I \ddot{\theta} + \Delta (GM_0 + GM_1 \cos(\omega t)) \theta = 0 \Rightarrow$$

$$\ddot{\theta} + \left(\omega_n^2 + C \omega_e^2 \cos(\omega_e t)\right) \theta = 0$$

where $\omega_n = \sqrt{\frac{gGM_0}{k}}$; $C = \frac{GM_1}{GM_0}$; $\Delta = mg$; $I = mk^2$; and $\omega_e$ = encounter wave freq.

By changing of variables ($\tau = \omega_e t$):

$$\ddot{\theta}(\tau) + \delta(1 + C \cos \tau) \theta(\tau) = 0$$

and $\delta = \frac{\omega_n^2}{\omega_e^2}$

This ordinary 2nd order differential equation where the restoring moment varies sinusoidally, is known as the Mathieu equation. This equation gives unbounded solution (i.e. it is unstable) when

$$\delta = \frac{\omega_n^2}{\omega_e^2} = \left(\frac{2n+1}{2}\right)^2 \quad n = 0,1,2,3,..$$

For the principle parametric roll resonance, $n=0$ i.e.

$$\omega_e = 2\omega_n \quad \frac{2\pi}{T_e} = 2 \times \frac{2\pi}{T_n} \Rightarrow T_n = 2T_e$$
Case (2) Rigid Body Translation or Rotation

In rigid body motion, all particles are in combined translation and/or rotation and there is no relative motion between particles; consequently, there are no strains or strain rates and the viscous term drops out of the N-S equation ($\mu \nabla^2 \mathbf{v} = 0$).

$$\nabla p = \rho (\underline{g} - \underline{a})$$

from which we see that $\nabla p$ acts in the direction of $(\underline{g} - \underline{a})$, and lines of constant pressure must be perpendicular to this direction (by definition, $\nabla f$ is perpendicular to $f = \text{constant}$).
The general case of rigid body translation/rotation is as shown. If the center of rotation is at O where \( \vec{V} = \vec{V}_0 \), the velocity of any arbitrary point P is:

\[
\vec{V} = \vec{V}_0 + \Omega \times \vec{r}_0
\]

where \( \Omega \) = the angular velocity vector

and the acceleration is:

\[
\frac{d\vec{V}}{dt} = a = \frac{d\vec{V}_0}{dt} + \Omega \times (\Omega \times \vec{r}_0) + \frac{d\Omega}{dt} \times \vec{r}_0
\]

1 = acceleration of O

2 = centripetal acceleration of P relative to O

3 = linear acceleration of P due to \( \Omega \)

Usually, all these terms are not present. In fact, fluids can rarely move in rigid body motion unless restrained by confining walls.
1.) **Uniform Linear Acceleration**

\[ a = \text{constant} \]

\[ = a_x \hat{i} + a_z \hat{k} \]  \hspace{1cm} \text{(2-1)}

\[ \nabla p = \rho (g - a) = \text{Constant} \]

\[ = -\rho \left[(g + a_z) \hat{k} + a_x \hat{i}\right] \]

\[ \frac{\partial p}{\partial x} = -\rho a_x \]

1. \( a_x < 0 \) \quad \text{p increase in } +x

2. \( a_x > 0 \) \quad \text{p decrease in } +x

\[ \frac{\partial p}{\partial z} = -\rho (g + a_z) \]

1. \( a_z > 0 \) \quad \text{p decrease in } +z

2. \( a_z < 0 \text{ and } |a_z| < g \) \quad \text{p decrease in } +z \text{ but slower than } g

3. \( a_z < 0 \text{ and } |a_z| > g \) \quad \text{p increase in } +z
unit vector in the direction of $\nabla p$:

$$\hat{s} = \frac{\nabla p}{|\nabla p|} = \frac{(g + a_z)k + a_x\hat{i}}{\left[(g + a_z)^2 + a_x^2\right]^{\frac{1}{2}}}$$

lines of constant pressure are perpendicular to $\nabla p$.

$$n = \hat{s} \times j = \frac{a_x k - (g + a_z)\hat{i}}{\left[a_x^2 + (g + a_z)^2\right]^{\frac{1}{2}}}$$

unit vector in direction of $p=constant$

angle between $n$ and $x$ axes:

$$\theta = \tan^{-1} \frac{a_x}{(g + a_z)}$$

In general the pressure variation with depth is greater than in ordinary hydrostatics; that is:

$$\frac{dp}{ds} = \nabla p \cdot \hat{s} = \rho \left[a_x^2 + (g + a_z)^2\right]^{\frac{1}{2}} \frac{1}{\rho G}$$

which is $> \rho g$

$$p = \rho G s + \text{constant}$$

$$= \rho G s \quad \text{gage pressure}$$
2). Rigid Body Rotation

Consider a cylindrical tank of liquid rotating at a constant rate \( \Omega = \Omega_k \):

\[
\nabla p = \rho (g - a)
\]

\[
a = \Omega \times (\Omega \times r_0)
\]

\[=-r\Omega^2 \hat{e}_r\]

\[
\nabla p = \rho (g - a) = -\rho g \hat{k} + \rho r\Omega^2 \hat{e}_r
\]

i.e.

\[
\frac{\partial p}{\partial r} = \rho r\Omega^2
\]

\[
\frac{\partial p}{\partial z} = -\rho g
\]

and

\[
p = \frac{\rho}{2} r^2 \Omega^2 + f(z) + c
\]

\[
p_z = f' = -\rho g
\]

\[
f(z) = -\rho gz + C
\]

\[
p = \frac{\rho}{2} r^2 \Omega^2 - \rho gz + \text{Constant}
\]

The constant is determined by specifying the pressure at one point; say, \( p = p_0 \) at \( (r,z) = (0,0) \).

\[
p = p_0 - \rho gz + \frac{\rho}{2} r^2 \Omega^2
\]

(Note: Pressure is linear in \( z \) and parabolic in \( r \))
Curves of constant pressure are given by:

\[ z = \frac{p_0 - p}{\rho g} + \frac{r^2 \Omega^2}{2g} = a + br^2 \]

which are paraboloids of revolution, concave upward, with their minimum points on the axis of rotation.

The unit vector in the direction of \( \nabla p \) is:

\[ \hat{s} = \frac{-\rho g \hat{k} + \rho r \Omega^2 \hat{e}_r}{\left[ (\rho g)^2 + (\rho r \Omega^2)^2 \right]^{1/2}} \]

\[ \tan \theta = \frac{dz}{dr} = -\frac{g}{r \Omega^2} \quad \text{slope of } \hat{s} \]

\[ -\frac{\Omega^2}{g} dz = \frac{dr}{r} \rightarrow -\frac{\Omega^2 z}{g} = \ln r \]

i.e. \( r = C_1 \exp \left( -\frac{\Omega^2 z}{g} \right) \) \( \text{equation of } \nabla p \text{ surfaces} \)

The position of the free surface is found, as it is for linear acceleration, by conserving the volume of fluid.
Case (3) Pressure Distribution in Irrotational Flow; Bernoulli Equation

Navier-Stokes:

\[ \rho \dot{a} = -\nabla (p) - \rho g \hat{k} + \mu \nabla^2 \text{V} = -\nabla (p + \gamma z) + \mu \nabla^2 \text{V} \]

\[ \rho \left[ \frac{\partial \text{V}}{\partial t} + \text{V} \cdot \nabla \text{V} \right] = -\nabla (p + \gamma z) + \mu \left[ \nabla (\text{V} \cdot \text{V}) - \nabla \times (\nabla \times \text{V}) \right] \]

Viscous term=0 for \( \rho = \)constant and \( \omega = 0 \), i.e., Potential flow solutions also solutions NS under such conditions!

1. Assuming inviscid flow: \( \mu = 0 \)

Using vector identity \( \text{V} \cdot \nabla \text{V} = \frac{1}{2} \nabla \text{V} \cdot \text{V} - \text{V} \times (\nabla \times \text{V}) \)

\[ \rho \left[ \frac{\partial \text{V}}{\partial t} + \left( \frac{1}{2} \nabla \text{V} \cdot \text{V} - \text{V} \times (\nabla \times \text{V}) \right) \right] = -\nabla (p + \gamma z) \text{ Euler Equation} \]

2. Assuming inviscid and incompressible flow: \( \mu = 0 \), \( \rho = \)constant

\[ \frac{\partial \text{V}}{\partial t} + \nabla \left[ \frac{\text{V}^2}{2} + \frac{p}{\rho} + g z \right] = \text{V} \times \omega \quad \text{V}^2 = \text{V} \cdot \text{V} \quad (\omega \neq 0) \]
3. Assuming inviscid, incompressible and steady flow: \( \mu = 0, \rho = \text{constant}, \frac{\partial}{\partial t} = 0 \)

\[ \nabla B = \mathbf{V} \times \boldsymbol{\omega} \]

\[ B = \frac{V^2}{2} + \frac{p}{\rho} + gz \]

Consider:

\( \nabla B \) perpendicular \( B = \text{constant} \)

\( \mathbf{V} \times \boldsymbol{\omega} \) perpendicular \( \mathbf{V} \) and \( \boldsymbol{\omega} \)

Therefore, \( B = \text{constant} \) contains streamlines and vortex lines:
\[ \hat{e}_s \cdot \nabla B = \frac{\partial B}{\partial s} = 0 \]
\[ \hat{e}_r \cdot \nabla B = 0 \]
\[ B = \frac{V^2}{2} + \frac{p}{\rho} + gz = \text{constant along streamlines and vortex lines.} \]

4. Assuming inviscid, incompressible, steady and irrotational flow: \( \mu = 0, \rho = \text{constant}, \frac{\partial}{\partial t} = 0, \omega = 0 \)
\[ \nabla B = 0 \quad B = \text{constant (everywhere same constant)} \]

5. Unsteady inviscid, incompressible, and irrotational flow:
\( \mu = 0, \rho = \text{constant}, \omega = 0 \)
\[ \overline{V} = \nabla \phi \]
\[ V^2 = \nabla \phi \cdot \nabla \phi \]
\[ \nabla \left[ \frac{\partial \phi}{\partial t} + \frac{\nabla \phi \cdot \nabla \phi}{2} + \frac{p}{\rho} + gz \right] = 0 \]
\[ \frac{\partial \phi}{\partial t} + \frac{\nabla \phi \cdot \nabla \phi}{2} + \frac{p}{\rho} + gz = B(t) \]
\[ B(t) = \text{time dependent constant} \]
Alternate derivation using stream line coordinates:

\[ \mathbf{V} = v_s (s, t) \hat{e}_s + v_n \hat{e}_n = v_s (s, t) \hat{e}_s \]

\[ \nabla = \hat{e}_s \frac{\partial}{\partial s} + \hat{e}_n \frac{\partial}{\partial n} \]

\[ a = \frac{DV}{Dt} = \frac{\partial V}{\partial t} + V \cdot \nabla V = \frac{\partial V}{\partial t} + v_s \frac{\partial V}{\partial s} = \left[ \frac{\partial v_s}{\partial t} \hat{e}_s + v_s \frac{\partial \hat{e}_s}{\partial t} \right] + v_s \left[ \frac{\partial v_s}{\partial s} \hat{e}_s + v_s \frac{\partial \hat{e}_s}{\partial s} \right] \]

\[ a = \left[ \frac{\partial v_s}{\partial t} + v_s \frac{\partial v_s}{\partial s} \right] \hat{e}_s + \left[ -v_s \frac{\partial \theta}{\partial t} - \frac{v_s^2}{R} \right] \hat{e}_n \]

\[ \frac{\partial v_s}{\partial t} \text{ = local } a_s \text{ in direction of flow} \]

\[ \frac{\partial v_n}{\partial t} = -v_s \frac{\partial \theta}{\partial t} \text{ = local } a_n \text{ normal to flow} \]

\[ v_s \frac{\partial v_s}{\partial s} \text{ = convective } a_s \text{ due to convergence/divergence of streamlines} \]

\[ -\frac{v_s^2}{R} \text{ = normal } a_n \text{ due to streamline curvature} \]
Euler Equation

\[ \rho a = -\nabla (p + \gamma z) \]

Steady flow s equation:

\[ \rho v_s \frac{\partial v_s}{\partial s} = -\frac{\partial}{\partial s} (p + \gamma z) \]

\[ \frac{\partial}{\partial s} \left( \frac{v_s^2}{2} + \frac{p}{\rho} + gz \right) = 0 \]

i.e., B=constant along streamline

Steady flow n equation:

\[ -\rho \frac{\partial v_n^2}{R} = -\frac{\partial}{\partial n} (p + \gamma z) \]

\[ -\int \frac{v_n^2}{R} dn + \frac{p}{\rho} + gz = \text{constant across streamline} \]

Larger speed/density or smaller R require larger pressure gradient or elevation gradient normal to streamline.
Flow Patterns: Streamlines, Streaklines, Pathlines

1) A streamline is a line everywhere tangent to the velocity vector at a given instant.
2) A pathline is the actual path traveled by a given fluid particle.

3) A streakline is the locus of particles which have earlier passed through a particular point.

Note:
1. For steady flow, all 3 coincide.
2. For unsteady flow, $\psi(t)$ pattern changes with time, whereas pathlines and streaklines are generated as the passage of time.

**Streamline**

By definition we must have $\mathbf{V} \times dr = 0$ which upon expansion yields the equation of the streamlines for a given time $t = t_1$

$$\frac{dx}{u} = \frac{dy}{v} = \frac{dz}{w} = ds \quad s = \text{integration parameter}$$

So if $(u,v,w)$ known, integrate with respect to $s$ for $t = t_1$ with IC $(x_0,y_0,z_0,t_0)$ at $s = 0$ and then eliminate $s$.

**Pathline**

The pathline is defined by integration of the relationship between velocity and displacement.

$$\frac{dx}{dt} = u \quad \frac{dy}{dt} = v \quad \frac{dz}{dt} = w$$

Integrate $u,v,w$ with respect to $t$ using IC $(x_0,y_0,z_0,t_0)$ then eliminate $t$.

**Streakline**
To find the streakline, use the integrated result for the pathline retaining time as a parameter. Now, find the integration constant which causes the pathline to pass through \((x_0, y_0, z_0)\) for a sequence of times \(\xi < t\). Then eliminate \(\xi\).

**Example:** an idealized velocity distribution is given by:

\[
\begin{align*}
  u &= \frac{x}{1+t} \\
  v &= \frac{y}{1+2t} \\
  w &= 0
\end{align*}
\]

calculate and plot: 1) the streamlines 2) the pathlines 3) the streaklines which pass through \((x_0, y_0, z_0)\) at \(t=0\).

1.) First, note that since \(w=0\) there is no motion in the \(z\) direction and the flow is 2-D

\[
\begin{align*}
  \frac{dx}{ds} &= \frac{x}{1+t} & \frac{dy}{ds} &= \frac{y}{1+2t} \\
  x &= C_1 \exp\left(\frac{s}{1+t}\right) & y &= C_2 \exp\left(\frac{s}{1+2t}\right) \\
  s &= 0 \quad \text{at} \quad (x_0, y_0): \quad C_1 = x_0 \quad C_2 = y_0
\end{align*}
\]

and eliminating \(s\)
\[ s = (1 + t) \ln \frac{x}{x_0} = (1 + 2t) \ln \frac{y}{y_0} \]

\[ y = y_0 \left( \frac{x}{x_0} \right)^n \]

where \( n = \frac{1 + t}{1 + 2t} \)

This is the equation of the streamlines which pass through \((x_0, y_0)\) for all times \(t\).

2.) To find the pathlines we integrate

\[ \frac{dx}{dt} = \frac{x}{1 + t} \quad \frac{dy}{dt} = \frac{y}{1 + 2t} \]

\[ x = C_1 (1 + t) \quad y = C_2 (1 + 2t)^{1/2} \]

\[ t = 0 \quad (x, y) = (x_0, y_0) : \quad C_1 = x_0 \quad C_2 = y_0 \]

now eliminate \(t\) between the equations for \((x, y)\)

\[ y = y_0 \left[ 1 + 2 \left( \frac{x}{x_0} - 1 \right) \right]^{1/2} \]

This is the pathline through \((x_0, y_0)\) at \(t=0\) and does not coincide with the streamline at \(t=0\).
3.) To find the streakline, we use the pathline equations to find the family of particles that have passed through the point \((x_0, y_0)\) for all times \(\xi < t\).

\[
x = C_1 (1 + t) \quad y = C_2 (1 + 2t)^{\frac{1}{2}}
\]

\[
C_1 = \frac{x_0}{1 + \xi} \quad C_2 = \frac{y_0}{(1 + 2\xi)^{\frac{1}{3}}}
\]

\[
\xi = (1 + t) \frac{x_0}{x} - 1 = \frac{1}{2} \left[ (1 + 2t) \left( \frac{y_0}{y} \right)^2 - 1 \right]
\]

\[
\left( \frac{y}{y_0} \right)^2 = \frac{1 + 2t}{1 + 2[(1 + t) \left( \frac{x_0}{x} \right) - 1]}
\]

\[t = 0: \quad \frac{y}{y_0} = \left[ 1 + 2 \left( \frac{x_0}{x} - 1 \right) \right]^{\frac{1}{2}}\]
The Stream Function

Powerful tool for 2-D flow in which $\mathbf{V}$ is obtained by differentiation of a scalar $\psi$ which automatically satisfies the continuity equation.

Note for 2D flow

$$\nabla \times \mathbf{V} = \left( \frac{\partial w}{\partial y} - \frac{\partial v}{\partial z}, \frac{\partial w}{\partial z} - \frac{\partial u}{\partial z}, \frac{\partial w}{\partial x} - \frac{\partial u}{\partial y} \right) = (0, 0, \omega_z)$$

Continuity:

$$u_x + v_y = 0$$

say: $u = \psi_y$ and $v = -\psi_x$

then: $\frac{\partial}{\partial x}(\psi_y) + \frac{\partial}{\partial y}(-\psi_x) = \psi_{yx} - \psi_{xy} = 0$ by definition!

$$\mathbf{V} = \psi_y \hat{i} - \psi_x \hat{j}$$

curl$\mathbf{V} = \hat{k} \omega_z = -\hat{k} \nabla^2 \psi$  \hspace{1cm} ($\omega_z = v_x - u_y = -\psi_{xx} - \psi_{yy} = -\nabla^2 \psi$)

curl($\rho \frac{D\mathbf{V}}{Dt}$) = $-\nabla(p + \gamma z) + \mu \nabla^2 \mathbf{V}$

$$\rho \text{curl}(\nabla \cdot \nabla \mathbf{V}) = \mu \nabla^2 \text{curl} \mathbf{V} \hspace{1cm} \text{Steady constant property flow}$$

$$\rho(u \frac{\partial}{\partial x} + v \frac{\partial}{\partial y})(-\hat{k} \nabla^2 \psi) = \mu \nabla^2 (-\hat{k} \nabla^2 \psi)$$

$$\rho \left[ \frac{\partial \psi}{\partial y} \frac{\partial}{\partial x} (\nabla^2 \psi) - \frac{\partial \psi}{\partial x} \frac{\partial}{\partial y} (\nabla^2 \psi) \right] = \mu \nabla^4 \psi \hspace{1cm} \text{single scalar equation, but 4th order!}$$

boundary conditions (4 required):

at infinity: $u = \psi_y = U_\infty, \quad v = -\psi_x = 0$

on body: $u = v = 0 = \psi_y = -\psi_x$
**Irrotational Flow**

\[ \nabla^2 \psi = 0 \quad \text{2nd order linear Laplace equation} \]

on \( S_\infty \) : \( \psi = U_\infty y + \text{const.} \)

on \( S_B \) : \( \psi = \text{const.} \)

\[ u = \psi_y = \phi_x \]
\[ v = -\psi_x = \phi_y \]

\( \Psi \) and \( \phi \) are orthogonal.

\[ d\phi = \phi_x dx + \phi_y dy = udx + vdy \]
\[ d\psi = \psi_x dx + \psi_y dy = -vdx + udy \]

i.e. \( \left. \frac{dy}{dx} \phi = \text{const} \right| u = \frac{-1}{v} \left. \frac{dy}{dx} \psi = \text{const} \right| 

Flow net (streamlines) and (equipotential lines)
Geometric Interpretation of $\psi$

Besides its importance mathematically $\psi$ also has important geometric significance.

$\psi = \text{constant} = \text{streamline}$

Recall definition of a streamline:

$$\nabla \times \mathbf{dr} = 0 \quad \mathbf{dr} = dx\hat{i} + dy\hat{j}$$

$$\frac{dx}{u} = \frac{dy}{v}$$

$udy - vdx = 0$

compare with $d\psi = \psi_x dx + \psi_y dy = -vdx + udy$

i.e. $d\psi = 0$ along a streamline

Or $\psi =$constant along a streamline and curves of constant $\psi$ are the flow streamlines. If we know $\psi(x, y)$ then we can plot $\psi =$ constant curves to show streamlines.
Physical Interpretation

\[ dQ = V \cdot n \, dA \]

\[ = (i \frac{\partial \psi}{\partial y} - j \frac{\partial \psi}{\partial x}) \left( \frac{dy}{ds} \hat{i} - \frac{dx}{ds} \hat{j} \right) ds \times 1 \]

\[ = \psi_y \, dy + \psi_x \, dx \]

\[ = d\psi \]

(note that \( \psi \) and \( Q \) have same dimensions: \( \text{m}^3/\text{s} \))

i.e. change in \( d\psi \) is volume flux and across streamline \( dQ = 0 \).

\[ Q_{1 \rightarrow 2} = \int_{1}^{2} V \cdot n \, dA = \int_{1}^{2} d\psi = \psi_{2} - \psi_{1} \]

Consider flow between two streamlines:
Incompressible Plane Flow in Polar Coordinates

\[ \frac{1}{r} \frac{\partial}{\partial r} (rv_r) + \frac{1}{r} \frac{\partial}{\partial \theta} (v_\theta) = 0 \]

or:

\[ \frac{\partial}{\partial r} (rv_r) + \frac{\partial}{\partial \theta} (v_\theta) = 0 \]

say:

\[ v_r = \frac{1}{r} \frac{\partial \psi}{\partial \theta}, \quad v_\theta = -\frac{\partial \psi}{\partial r} \]

then:

\[ \frac{\partial}{\partial r} \left( r \frac{1}{r} \frac{\partial \psi}{\partial \theta} \right) + \frac{\partial}{\partial \theta} \left( -\frac{\partial \psi}{\partial r} \right) = 0 \]

as before \( d\psi = 0 \) along a streamline and \( dQ = d\psi \)

volume flux = change in stream function

Incompressible axisymmetric flow

\[ \frac{1}{r} \frac{\partial}{\partial r} \left( rv_r \right) + \frac{\partial}{\partial z} \left( v_z \right) = 0 \]

say:

\[ v = -\frac{1}{r} \frac{\partial \psi}{\partial z}, \quad \frac{v_z}{z} = \frac{1}{r} \frac{\partial \psi}{\partial r} \]

then:

\[ \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{1}{r} \frac{\partial \psi}{\partial z} \right) + \frac{\partial}{\partial z} \left( \frac{1}{r} \frac{\partial \psi}{\partial r} \right) = 0 \]

as before \( d\psi = 0 \) along a streamline and \( dQ = d\psi \)
**Generalization**

Steady plane compressible flow:

\[
\frac{\partial}{\partial x} (\rho u) + \frac{\partial}{\partial y} (\rho v) = 0
\]

define: \( \rho u = \frac{\partial \psi}{\partial y} \quad \rho v = -\frac{\partial \psi}{\partial x} \quad \psi = \text{compressible flow stream function} \)

Alongside \( \psi \quad u\,dy - v\,dx = 0 \)

compare with \[
\frac{1}{\rho} \psi_y dy + \frac{1}{\rho} \psi_x dx = 0
\]

\[
d\psi = \psi_x dx + \psi_y dy \Rightarrow \frac{1}{\rho} (d\psi) = 0 \quad \text{i.e.} \quad d\psi = 0 \quad \text{and} \quad \psi = \text{constant is a streamline}
\]

Now:

\[
d\dot{m} = \rho (\nabla \cdot \vec{n}) dA = d\psi
\]

\[
\dot{m}_{1\rightarrow 2} = \int_{1}^{2} \rho (\nabla \cdot \vec{n}) dA = \psi_2 - \psi_1
\]

Change in \( \psi \) is equivalent to the mass flux.